

## Aberystwyth University

### Packing of Softly Repulsive Particles in a Spherical Box —A Generalised Thomson Problem

Mughal, Adil

*Published in:*  
Forma

*DOI:*  
[10.5047/forma.2014.003](https://doi.org/10.5047/forma.2014.003)

*Publication date:*  
2014

*Citation for published version (APA):*

Mughal, A. (2014). Packing of Softly Repulsive Particles in a Spherical Box —A Generalised Thomson Problem. *Forma*, 29(1), 13-19. <https://doi.org/10.5047/forma.2014.003>

#### General rights

Copyright and moral rights for the publications made accessible in the Aberystwyth Research Portal (the Institutional Repository) are retained by the authors and/or other copyright owners and it is a condition of accessing publications that users recognise and abide by the legal requirements associated with these rights.

- Users may download and print one copy of any publication from the Aberystwyth Research Portal for the purpose of private study or research.
- You may not further distribute the material or use it for any profit-making activity or commercial gain
- You may freely distribute the URL identifying the publication in the Aberystwyth Research Portal

#### Take down policy

If you believe that this document breaches copyright please contact us providing details, and we will remove access to the work immediately and investigate your claim.

tel: +44 1970 62 2400  
email: [is@aber.ac.uk](mailto:is@aber.ac.uk)

# Packing of Softly Repulsive Particles in a Spherical Box —A Generalised Thomson Problem

Adil Mughal

*Institute of Mathematics and Physics, Aberystwyth University, Penglais, Aberystwyth, Ceredigion, Wales, SY23 3BZ, U.K., and  
Theoretische Physik, Fried.-Alex.-Universität Erlangen-Nürnberg - Staudtstr. 7, 91058 Erlangen, Germany, EU  
E-mail address: adil.m.mughal@gmail.com*

(Received December 29, 2013; Accepted April 25, 2014)

We study the (near or close to) ground state distribution of  $N$  softly repelling particles trapped in the interior of a spherical box. The charges mutually interact via an inverse power law potential of the form  $1/r^\gamma$ . We study three regimes in which the charges form a single spherical shell at the edge of the box ( $\gamma = 1$ ), a series of concentric shells of increasing density ( $\gamma = 2$ ) and  $\gamma = 12$  for which the charges form shells with a more uniform charge distribution. We conduct numerical simulations for clusters containing up to 5000 charges and compare charge density across the system with continuum limit results. The agreement between numerical (discrete) results and the continuum limit is found to improve with increasing  $N$ . These findings are accompanied by a visual gallery of some of the low energy ground states found by simulated annealing.

**Key words:** Thomson Problem, Low Temperature Plasma, Packing in a Spherical Box

## 1. Introduction

The *generalised* Thomson problem, as dubbed by Bowick *et al.* (2002), is concerned with finding the minimal energy configuration of  $N$  point particles, that are confined to the surface of a sphere, and repel each other via an inverse power law potential  $\phi(\gamma) = 1/r^\gamma$ . For over a decade the study of such problems has yielded fundamental insights into crystallisation and order on curved surfaces. Applications include understanding virus morphology (Zandi and Reguera, 2005), self-assembly of colloids on emulsion droplets (Bausch *et al.*, 2003) and multi-electron bubbles in superfluid helium (Bowick *et al.*, 2006).

However, the original Thomson problem has its origins in the “plum” pudding model of the atom. The model, proposed by J. J. Thomson in 1904, postulates that the atom consists of classical electrons embedded in a neutralising droplet of positively charged fluid (Thomson, 1904). Although this pre-quantum era model of the atom is now obsolete, nevertheless, the Thompson problem and its variations continue to be of interest to modern science in areas such as packing problems (Aste and Weaire, 2008), for benchmarking various optimisation algorithms (Cecka *et al.*, 2013) and as means of efficiently discretising space for lattice simulations (Hüttig and Stemmer, 2008).

In this paper we return to the original spirit of the Thomson problem and conduct numerical simulations to find (near) ground states for clusters of particles, in a spherical box of radius  $R$ , which interact through a potential  $\phi(\gamma)$ .

Such inverse power law potentials have been widely studied (see Agrawal and Kofke, 1995; Prestipino *et al.*, 2005; Bowick *et al.*, 2006) and provide a continuous path be-

tween the hard sphere packing limit ( $\gamma \rightarrow \infty$ ) and the soft one-component plasma ( $\gamma = 1$ ). In practice, such a path can be realised using dilute solutions of colloidal particles with polymer chains grafted onto their surface. The effective pairwise interaction between colloidal particles can be tuned by changing the thickness of the polymer layer and modelled with an inverse-power-law potential of the type  $\phi(\gamma)$  (Likos, 2001; Prestipino *et al.*, 2005).

In contrast to bulk systems (where at equilibrium particles—interacting via an inverse power law potential—crystallise into bcc or fcc arrangements) we show that in strongly confined systems the morphology of the ground state is dominated by the shape of the confining box. The main finding is that whereas for  $\gamma = 1$  (Coulomb interaction), the particles are pinned to the surface, for higher values of  $\gamma$  bulk aggregation is observed, either in the form of concentric shells (for intermediate  $\gamma$  values) or in the form of a quasi-uniform fluid (for higher gamma values). By tuning the effective interaction range in colloidal systems (for example by changing the screening length (Yethiraj and van Blaaderen, 2003) similar transitions could be observed in strongly confined systems.

In the absence of a neutralising fluid, and in the limit of large  $N$ , the continuum limit ground state distribution of particles within the spherical box falls into three distinct regimes. (i) For  $\gamma \leq 1$  the repulsive interaction is strong enough to drive all the particles to the surface of the sphere (Levin and Arenzon, 2003), that is the problem of finding the ground state of  $N$  point-particles inside a sphere reduces to the problem studied by Bowick *et al.* (2002, 2006) and Bausch *et al.* (2003). (ii) While for  $1 < \gamma < 3$  this is no longer the case and charges are found in the interior of the sphere with a non-uniform radial density. Finally, (iii) for  $\gamma \geq 3$  the particles are expected to be distributed

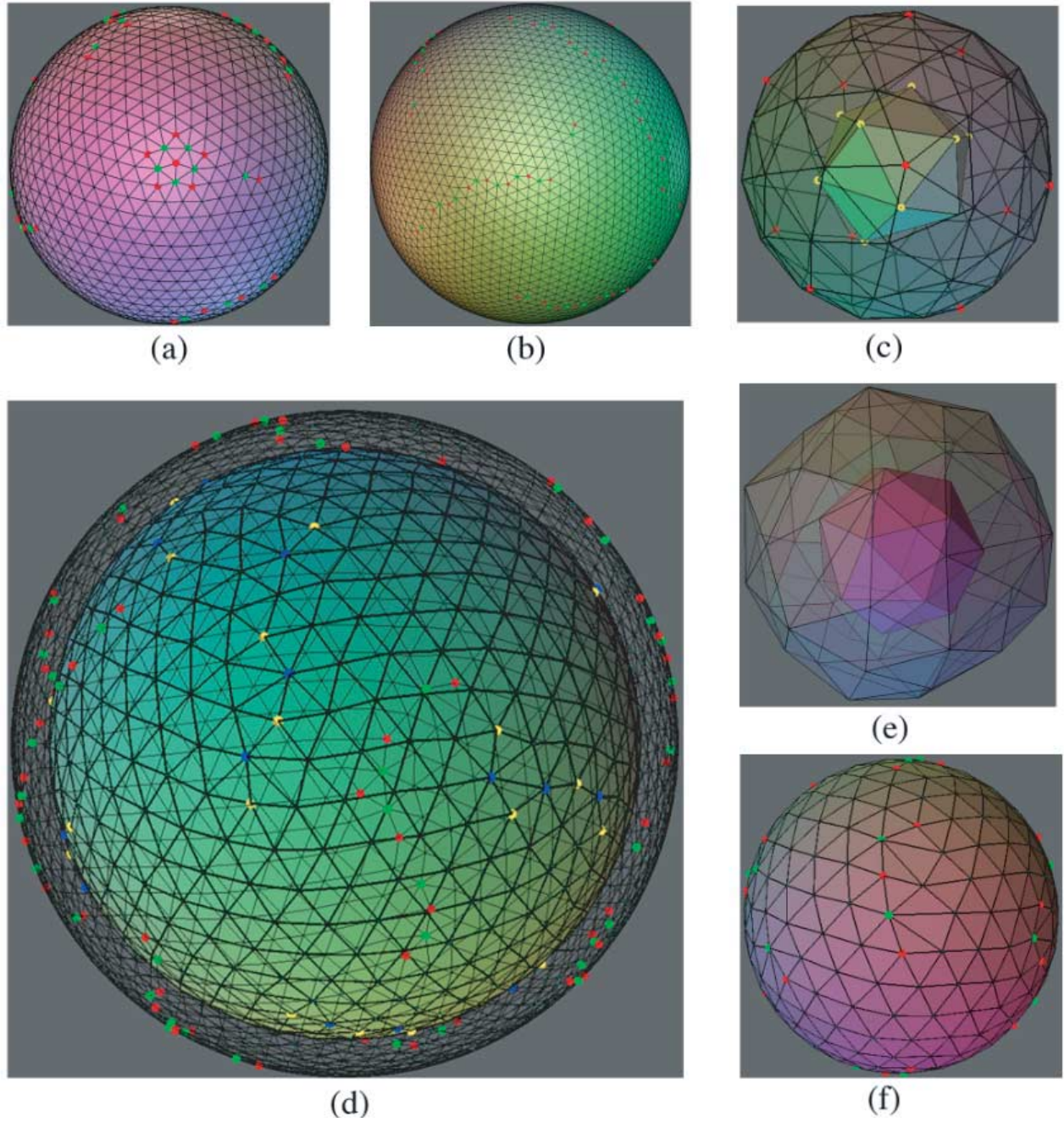


Fig. 1. Colour online. (a) Low energy metastable state of 2000 charges interacting via an inverse potential  $\gamma = 1$  (i.e. the Coulomb potential) showing an isolated rosette defect. (b) More typically however are grain boundary scars (of alternating positive and negative disclinations) as can be seen in this low energy metastable state of 5000 charges with  $\gamma = 1$ . Positive and negative disclinations in (a) and (b) are coloured red and green respectively. (c) A small cluster of 100 charges with  $\gamma = 2$ ; the particles are arranged into two distinct shells, an inner shell of 13 particles and an outer shell with 87 charges. (d) The two outermost shells for a system with 5000 charges and  $\gamma = 2$ . Positive and negative disclinations on the inner shells in (c) and (d) are coloured yellow and blue, respectively. (e) The two innermost shells in a cluster of 2000 charges with  $\gamma = 2$ . (f) Outer shell of a low energy state of 1000 charges interacting via an inverse potential with  $\gamma = 12$ .

throughout the sphere with uniform density (Hardin and Saff, 2004).

Although it is recognised that for  $\gamma > 1$  the presence of charges in the interior of the spherical box becomes energetically favourable (Levin and Arenzon, 2003), our simulations demonstrate that these internal charges also condense into spherical crystals (in which most charges have six nearest neighbours) giving rise to a series of concentric shells. Decreasing the range of interaction (by increasing  $\gamma$ ) drives

particles into the interior and leads to a higher occupancy of the inner shells. This is in contrast to similar shell like structures observed in spherical dusty plasma crystals (so called Yukawa balls), where decreasing the range of interaction (achieved by decreasing the screening length) drives particles towards the exterior and leads to a higher occupancy of the outer shells (Baumgartner *et al.*, 2009).

For  $\gamma \geq 3$  the continuum limit charge density is expected to be uniform. In finite sized clusters the charges

are found to be self-organise into multi-shell arrangements which could in principle be compared with similar carbon-based structures such as nested Fullerenes (Tománek *et al.*, 1993). Although the precise morphology of these latter systems depends on quantum chemistry, nevertheless the approach described here may prove useful in distinguishing between simple features, which are largely geometric in nature, and more complex properties arising from chemistry.

The ground state configuration of mutually repelling charges is of interest in the efficient discretisation of spaces as a framework for various numerical schemes (Hardin and Saff, 2004). These include statistical sampling, finite element tessellations, quadrature, interpolation and as starting points for Newton's method. In the case of a spherical geometry, in such problems the need can often arise for a greater degree of resolution either at the centre or towards the surface of the box (as for example in the modelling of seismic waves through the earth (Akcelik *et al.*, 2003; Grunberg *et al.*, 2004)). By fixing  $\gamma$  to be in the range  $1 < \gamma < 3$  the present study provides a simple method for generating a mesh with a varying radial node density.

The paper is organised as follows. We begin by detailing our numerical approach. We then elucidate the three regimes by considering three exemplary cases, these are:  $\gamma = 1, 2$  and  $12$ . In each instance we first state or derive the expected continuum limit charge distribution and then compare this with a series of finite sized clusters. In addition we also provide a brief pictorial gallery that is representative of some of the low energy structures found by our numerical simulations. We hope that these images will be of interest to both scientists (e.g. chemists studying multi-wall fullerenes) as well as nonscientists (e.g. designers working with buckyball architecture).

## 2. Numerical Approach

The energy of a cluster of  $N$  particles, interacting via a potential  $\phi(\gamma)$  and confined to the interior of a sphere of radius  $R$ , by a hard wall potential is given by,

$$E = \sum_i^N V(r_i) + \sum_{i < j}^N \frac{1}{|\mathbf{r}_i - \mathbf{r}_j|^\gamma}, \quad (1)$$

where

$$V(r_i) = \begin{cases} 0 & \text{for } r_i < R \\ \infty & \text{for } r_i \geq R \end{cases}$$

and  $\mathbf{r}_i = (r_i, \theta_i, \phi_i)$ . Finding the global minimum for a function such as Eq. (1) is a difficult task. The number of metastable states proliferate exponentially with  $N$ ; consequently the global minimum is obscured by a large number of local minima with energies close to that of the global minimum. There exist a number of heuristic methods for such problems. Although there is no guarantee of finding the global minimum, it is possible to find states close to it.

We found that the standard Metropolis simulated annealing algorithm to be more effective than a conjugate gradient algorithm. For a system with  $N$  particles the simulated annealing algorithm was run with typically  $N \times (5 \times 10^6)$  Monte Carlo steps. The temperature of the simulation was decreased linearly. The average displacement of the

charges at each temperature step was chosen by an automatic process to give an acceptance probability of  $0.5 \pm 0.01$ . Promising states were reheated and annealed repeatedly to iron out as many defects as possible. Finally the results were put through a conjugate gradient algorithm to remove any residual strains.

## 3. First Case: $\gamma = 1$

For  $\gamma \leq 1$  the inter particle repulsion is strong enough to drive all the charges to the edge of the spherical box. The Coulomb ( $\gamma = 1$ ) case is merely a reflection of the familiar result from electrostatics that, under static conditions, the charge density inside a conductor is always zero (Hardin and Saff, 2004). Thus we expect the charges to be located on the surface of the sphere and for the charge density, in the continuum limit, to be described by a delta function of the form,

$$\rho(r) = \frac{N}{4\pi R^2} \delta(r - R). \quad (2)$$

Setting  $\gamma = 1$  in Eq. (1), we readily find an exact correspondence between Eq. (2) and our numerical results for  $N = 1000, 2000$  (Fig. 1(a)) and  $5000$  (Fig. 1(b)). These show that *all* the charges are indeed located in a single shell at the edge of the system (no evidence of buckling was found).

Using the Delaunay triangulation package Qhull we identify the number of nearest neighbours for each particle. Particles with five/seven nearest neighbours are coloured red/green, while particles with six neighbours are not highlighted.

Euler's theorem stipulates that such spherical crystal cannot consist entirely of six coordinated particles but must also include a minimum of twelve pentagonal sites. Such points with an anomalous coordination are topological defects known as disclinations. Using Euler's theorem we can assign a topological charge to each disclination, the sign and magnitude of the charge depends on how much the coordination number differs from 6. Thus, a pentagon has a topological charge  $+1$  while a square has a charge of  $+2$ . Similarly a heptagon has a topological charge of  $-1$  while an octagon has  $-2$ . Obviously a hexagon is topologically neutral. The total topological charge for any spherical cluster is conserved and must always be equal to  $+12$  (Bausch *et al.*, 2003).

Such disclinations induce an enormous elastic strain in the lattice which can be reduced by arranging them symmetrically over the surface of the sphere. In addition the lattice may also include dislocations (tightly bound five-seven coordinated disclination pairs). Unlike disclinations, the number of dislocations is not fixed by topology and are only present if it is energetically favourable.

Large clusters (those with more than 520 particles (Wales *et al.*, 2009)) always contain dislocations. Typically, such dislocations condense around disclinations to form extended grain boundary "scars" of alternating positive-negative disclinations, as can be seen in the case for  $N = 5000$  (see Fig. 1(b)). The net topological charge of these scars is  $+1$ . More exotic disclination structures, such as rosettes are also possible. An example is shown in the low

energy state of a cluster of 2000 charges, see Fig. 1(a). Such defects have a net topological charge of +1 and consist of a central positive disclination surrounded by five negative disclinations alternating with five positive disclinations. Rosette arrangements have been observed in both spherical (Wales *et al.*, 2009) and flat crystals (Radzvilavicius and Anisimovas, 2011).

#### 4. Second Case: $\gamma = 2$

In the range  $1 < \gamma < 3$  the charge density is expected to be intermediate between being entirely concentrated at the boundary and being uniform throughout the sphere. We concentrate on  $\gamma = 2$  since in this case it is possible to derive the charge density in closed form.

For  $\gamma = 2$  the energy given by Eq. (1) can be approximated by the integrals over the sphere  $r \leq R$ ,

$$E = \frac{1}{2} \int d^3 r' \int d^3 r \frac{\rho(\mathbf{r}')\rho(\mathbf{r})}{|\mathbf{r} - \mathbf{r}'|^2}. \quad (3)$$

The continuum approximation treats the density  $\rho(\mathbf{r})$  as a smooth function rather than the sum of delta functions,

$$\rho(\mathbf{r}) = \sum_{i=1}^N \delta(\mathbf{r} - \mathbf{r}_i), \quad (4)$$

where  $\mathbf{r}_i$  is the position of the  $i^{th}$  charge. One then minimizes the energy of the cluster, with respect to the smooth function  $\rho(\mathbf{r})$ , subject to the constraint that the number of particles,

$$N = \int d^3 r \rho(\mathbf{r}), \quad (5)$$

is constant. Introducing the Lagrange multiplier  $\mu$  the constrained equation is,

$$E = \int \rho(\mathbf{r}') \left[ \frac{1}{2} \int d^3 r \frac{\rho(\mathbf{r})}{|\mathbf{r} - \mathbf{r}'|^2} - \mu \right] d^3 r'. \quad (6)$$

A variation in the energy is given by,

$$\delta E = E[\rho(\mathbf{r}) + \delta\rho(\mathbf{r})] - E[\rho(\mathbf{r})], \quad (7)$$

where  $\delta\rho(\mathbf{r})$  represents a small change in the charge density. Keeping only terms up to first order, Eq. (7) gives,

$$\delta E = \int \delta\rho(\mathbf{r}') \left[ \int d^3 r \frac{\rho(\mathbf{r})}{|\mathbf{r} - \mathbf{r}'|^2} - \mu \right] d^3 r', \quad (8)$$

where to make the functional derivative stationary we require that,

$$\mu = \int d^3 r' \frac{\rho(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|^2}. \quad (9)$$

Assuming spherical symmetry and expanding the denominator in the above equation as a power series, in terms of Legendre polynomials, yields,

$$\mu = 2\pi \int_0^R \rho(r') r'^2 dr' \int_0^\pi d\theta \sin\theta \left[ \frac{1}{r'^2} \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \left( \frac{r_{<}}{r_{>}} \right)^{n+m} P_n(\cos\theta) P_m(\cos\theta) \right] \quad (10)$$

where,

$$\cos\theta = \frac{\mathbf{r} \cdot \mathbf{r}'}{|\mathbf{r}||\mathbf{r}'|}, \quad (11)$$

and to ensure convergence  $r_{>}$  is the greater of  $r$  and  $r'$ . Upon making the substitution  $x = \cos\theta$  we can write Eq. (10) as,

$$\frac{\mu}{2\pi} = \sum_{n=0}^{\infty} \sum_{m=0}^{\infty} \int_0^R dr' \rho(r') \frac{r'^2}{r_{>}^2} \left( \frac{r_{<}}{r_{>}} \right)^{m+n} \int_{-1}^1 dx P_n(x) P_m(x), \quad (12)$$

and upon using the orthogonality conditions of Legendre polynomials we have,

$$\frac{\mu}{2\pi} = \sum_{n=0}^{\infty} \frac{4\pi}{2n+1} \int_0^R dr' \rho(r') \frac{r'^2}{r_{>}^2} \left( \frac{r_{<}}{r_{>}} \right)^{2n}. \quad (13)$$

Equation (13) can be split into two parts  $r > r'$  and  $r < r'$  giving,

$$\frac{\mu}{2\pi} = \sum_{n=0}^{\infty} \frac{4\pi}{2n+1} \left[ \int_0^r dr' \rho(r') \left( \frac{r'}{r} \right)^{2n+2} + \int_r^R dr' \rho(r') \left( \frac{r}{r'} \right)^{2n} \right], \quad (14)$$

and writing the power series in Eq. (14) in closed form we have,

$$\frac{\mu r}{2\pi} = \int_0^r dr' g(r') r' \ln \left( \frac{r+r'}{r-r'} \right) - \int_r^R dr' g(r') r' \ln \left( \frac{r+r'}{r-r'} \right), \quad (15)$$

where  $g(r') = r' \rho(r')$ . Differentiating both sides of Eq. (15) with respect to  $r$  yields,

$$\frac{\mu}{2\pi} = \int_0^R dr' g(r') \frac{1}{r+r'} - \int_0^r dr' g(r') \frac{1}{r-r'} - \int_r^R dr' g(r') \frac{1}{r'-r}. \quad (16)$$

Treating the last two integrals in Eq. (16) as a Cauchy principle value integral we finally have,

$$\frac{\mu}{2\pi} = \int_0^R dr' g(r') \frac{2r'}{r'^2 - r^2}. \quad (17)$$

Equation (17) is a singular integral equation of the first kind which can be solved to give (Polyanin and Manzhirov, 2008),

$$\rho(r) = \frac{\mu}{2\pi^2} \left( \frac{1}{R^2 - r^2} \right)^{1/2}. \quad (18)$$

To find the Lagrange multiplier we substitute Eq. (18) into Eq. (5) and solve for  $\mu$ , finally we have that the particle density is,

$$\rho(r) = \frac{N}{\pi^2} \frac{1}{R^3 \left( 1 - \frac{r^2}{R^2} \right)^{1/2}}. \quad (19)$$



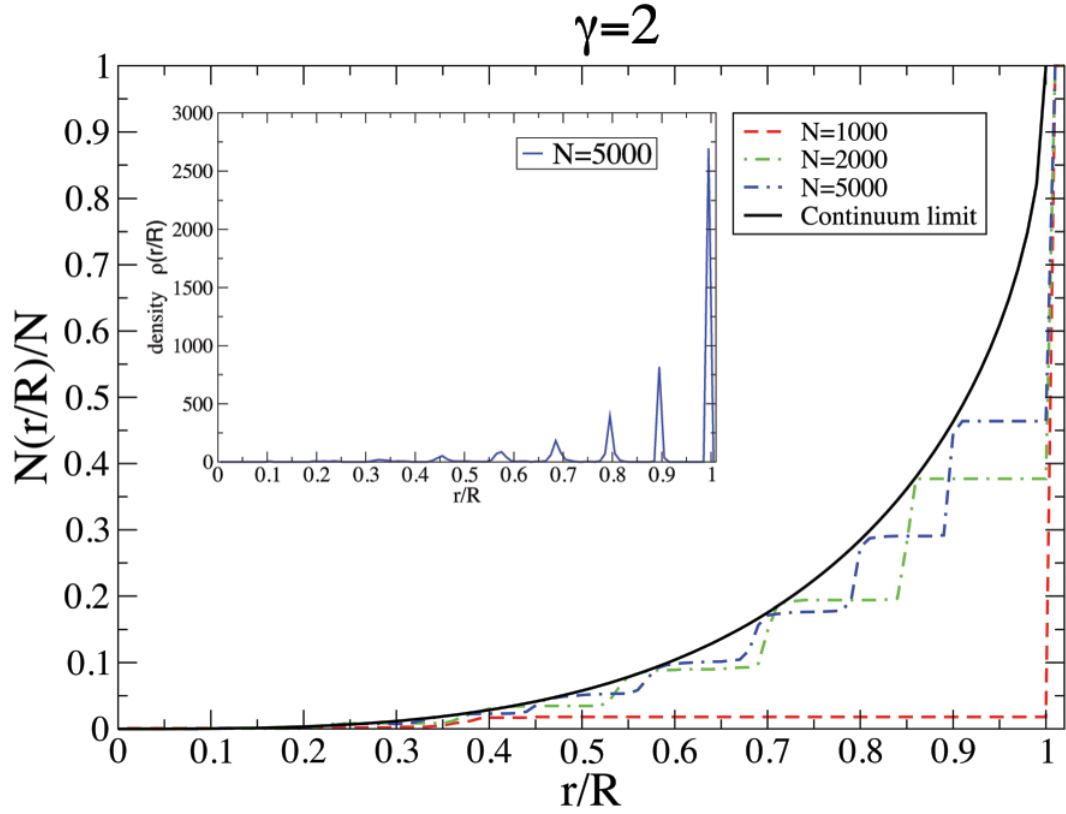


Fig. 2. Histograms of the fraction of the total charge of the system within a given radius  $r/R$  for  $\gamma = 2$ . Results for clusters containing 1000, 2000, and 5000 charges are coloured red, green and blue respectively. The continuum limit result is given by the black dashed line. The inset shows the charge density for the system with 5000 particles.

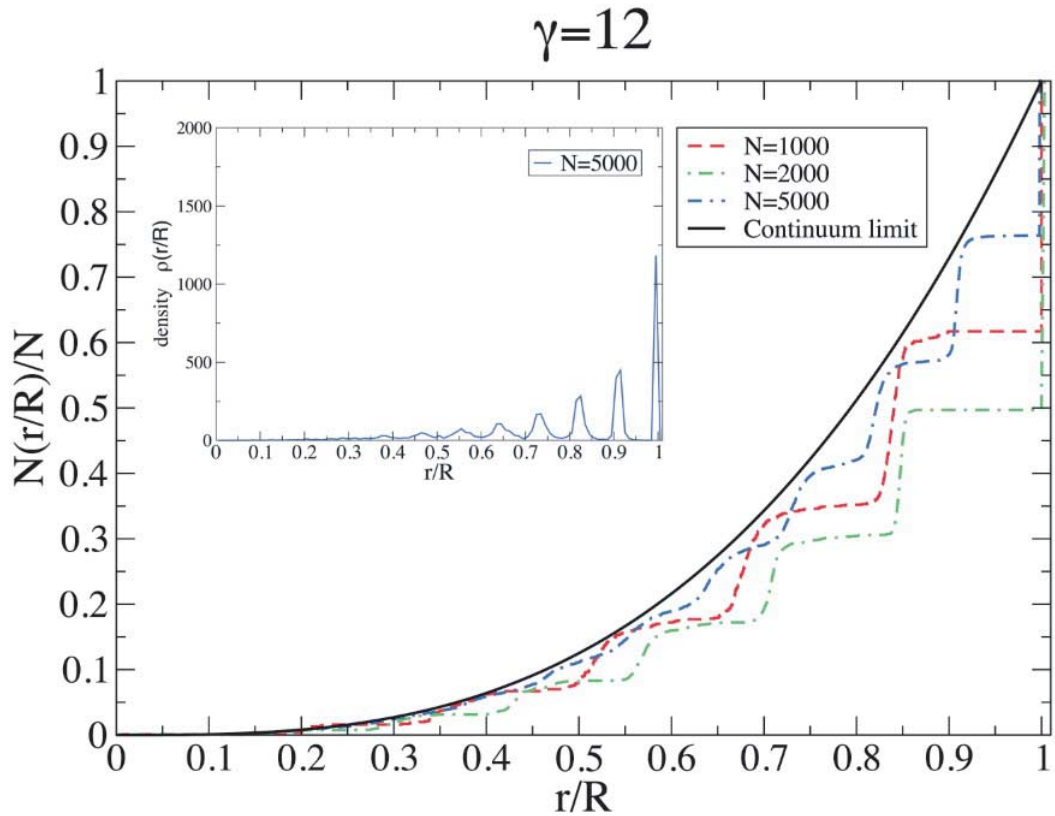


Fig. 3. Histograms of the fraction of the total charge of the system within a given radius  $r/R$  for  $\gamma = 12$ . Results for clusters containing 1000, 2000, and 5000 charges are coloured red, green and blue respectively. The continuum limit result is given by the black dashed line. The inset shows the charge density for the system with 5000 particles.

Thus in the continuum limit, unlike the case for  $\gamma = 1$ , not all of the charge is found at the edge of the system. The fraction of charges within a fractional distance  $r/R$  from the centre of the bounding sphere is given by integrating Eq. (19), giving

$$\frac{N\left(\frac{r}{R}\right)}{N} = \frac{2}{\pi} \left( \sin^{-1}\left(\frac{r}{R}\right) - \frac{r}{R} \sqrt{1 - \left(\frac{r}{R}\right)^2} \right). \quad (20)$$

Equation (20) can be compared with numerical results for low energy clusters with  $N = 1000, 2000$  and  $5000$  charges, see Fig. 2. As expected, the agreement between numerical and analytical results improves with increasing  $N$ .

In the case of such finite sized clusters the distribution shown in Fig. 2 displays a step like behaviour, particularly close to the edge of the system, which indicates that the charges form a series of concentric shells around the centre of the spherical box. This corresponds to a series of sharp well defined peaks of increasing density as shown in the inset in Fig. 2 for the  $N = 5000$  system.

The morphology of small clusters ( $N \lesssim 500$ ) in particular is dominated by the spherical hard wall boundary. In small systems all of the charges are arranged into well defined shells, where Euler's theorem holds individually for each shell. An example of a system with 100 charges is shown in Fig. 1(c).

For larger systems this is only true close to the edge of the system where the spherical boundary forces the charges to be concentrated into shells, corresponding to sharp peaks in density. Figure 1(d) shows the two outermost layers for  $N = 5000$ . In both shells twelve grain boundary scars comprised of alternating positive and negative disclinations can be observed. However, towards the centre of the system the influence of the spherical boundary diminishes and the peaks in density become broader and less well defined. Consequentially, the shells in the inner region are irregularly shaped and it becomes difficult to uniquely identify distinct shells, see for example Fig. 1(e) which shows the two innermost shells for  $N = 2000$ . This task is made more difficult by the presence of numerous isolated inter-shell charges in the interior region, i.e. charges which are found between adjacent shells and cannot be said to belong to either shell. Furthermore, the system for  $N = 1000$  appears to be characteristically different from that of  $N = 2000$  and  $N = 5000$ . The latter two are comprised of numerous concentric shells while the former is not. Clearly there is a critical value beyond  $N = 1000$  at which the internal shells start to fill up rapidly. A theoretical understanding of where this critical point is located may be an interesting future problem.

### 5. Third Case: $\gamma = 12$

For  $\gamma \geq 3$  the charge density in the continuum limit is expected to be uniform (Hardin and Saff, 2004). We focus on the case of  $\gamma = 12$  which is the repulsive part of the familiar Lennard-Jones potential. Once again, we plot the fraction of charges  $N(r/R)/N$ , within a fractional radius  $r/R$  and compare with the continuum limit result, see

Fig. 3. We find an improving agreement between analytical and numerical results.

However, it is clear that in finite clusters there exists a systematic deviation from the expected uniform charge density, as seen by the fact that the peaks in density are much higher towards the edge of the system, see inset in Fig. 3 for  $N = 5000$ . Or to put it another way: the mass fraction in inner region is always slightly smaller than the theoretical curve. Such deviations are to be expected (see Mughal and Moore, 2007) for a comparative example) and a closer match to the charge distribution in finite clusters can be provided by higher order correction to the density that takes into account the shell like structure of the system close to the spherical boundary. It is possible that further annealing may yield a more uniform charge distribution but this has so far not proved to be the case, despite extensive numerical efforts.

Again, near the spherical boundary the charges form a series of concentric spherical crystals, as an example the outermost layer in a system of 1000 charges is shown in Fig. 1(f). However, towards the centre these shells are increasingly deformed, and it becomes difficult to uniquely identify separate shells (inter shell charges are also observed in the interior region).

Similar arrangements of nested shells, of uniform density, are also *expected* to be observed in spherical clusters made up of single-species of charged particle in a three-dimensional confining potential. In such systems the transition to a bcc lattice is predicted to occur when the number of charges in the system increases beyond 10000 (Totsuji *et al.*, 2002). Similarly, in the case of inverse power law interacting charges in a spherical box the emergence of a crystalline interior for  $\gamma \geq 3$  may only emerge when the number of charges is sufficiently great enough to screen the interior of the system from the presence of the spherical boundary.

### 6. Conclusions

We studied a variant of the Thomson problem in which we seek the minimal energy arrangement of generalised charges (interacting via an inverse power law potential with exponent  $\gamma$ ) in a spherical box. We find that the structure of the cluster, especially whether the particles are localised to the surface of the sphere, depends on the power law tail of the interaction. Increasing the value of  $\gamma$  drives charges from the edge of the box into the interior. We find that in finite sized clusters the charges close to the edge of the spherical box are arranged into a series of concentric shells, within which the charges form well defined spherical crystals. The presence of these shell like structures creates significant deviation from the expected mass fraction results. However, towards the interior of the box the influence of the boundary is diminished and charges are arranged into less well defined configurations.

**Acknowledgments.** AM acknowledges useful discussions with Mike Moore.

### References

Agarwal, R. and Kofke, D. A. (1995) *Phys. Rev. Lett.*, **74**, 122–125.

- Akcelik, V., Bielak, J., Biros, G., Epanomeritakis, I., Fernandez, A., Ghattas, O., Kim, E. J., Lopez, J., O'Hallaron, D., Tu, T. and Urbanic, J. (2003) High resolution forward and inverse earthquake modeling on terascale computers, in *Supercomputing, 2003 ACM/IEEE Conference*, pp. 52–52, IEEE.
- Aste, T. and Weaire, D. (2008) *The Pursuit of Perfect Packing*, Taylor and Francis Group.
- Baumgartner, H., Block, D. and Bonitz, M. (2009) *Cont. to Plasma Phys.*, **49**, 281.
- Bausch, A., Bowick, M., Cacciuto, A., Dinsmore, A., Hsu, M., Nelson, D., Nikolaides, M., Travesset, A. and Weitz, D. (2003) *Science*, **299**, 1716.
- Bowick, M., Cacciuto, A., Nelson, D. R. and Travesset, A. (2002) *Phys. Rev. Lett.*, **89**, 185502.
- Bowick, M., Cacciuto, A., Nelson, D. R. and Travesset, A. (2006) *Phys. Rev. B*, **73**, 024115.
- Cecka, C., Middleton, A. and Bowick, M. (2013) Thomson problem database and applet, May 2013. Available at: <http://thomson.phy.syr.edu/>
- Grunberg, M., Genaud, S. and Mongenet, C. (2004) *The Journal of Supercomputing*, **29**, 27.
- Hardin, D. P. and Saff, E. B. (2004) *Not. Am. Math. Soc.*, **51**, 1186–1194.
- Hüttig, C. and Stemmer, K. (2008) *Geochem., Geophys., Geosys.*, **9**.
- Levin, Y. and Arenzon, J. J. (2003) *Europhys. Lett.*, **63**, 415–418.
- Likos, C. N. (2001) *Phys. Rep.*, **348**, 267.
- Mughal, A. and Moore, M. A. (2007) *Phys. Rev. E*, **76**, 011606.
- Polyanin, A. D. and Manzhirov, A. V. (2008) *Handbook of Integral Equations*, Taylor and Francis Group.
- Prestipino, S., Saija, F. and Giaquinta, P. V. (2005) *J. Chem. Phys.*, **123**, 144110.
- Radzvilavicius, A. and Anisimovas, E. (2011) *J. Phys.: Condens. Matter*, **23**, 385301.
- Thomson, J. (1904) *Phil. Mag.*, **73**, 237.
- Tománek, D., Zhong, W., Krastev, E. *et al.* (1993) *Phys. Rev. B*, **48**, 15461.
- Totsuji, H., Kishimoto, T., Totsuji, C. and Tsuruta, K. (2002) *Phys. Rev. Lett.*, **88**, 125002.
- Wales, D. J., McKay, H. and Altschuler, E. L. (2009) *Phys. Rev. B*, **79**, 224115.
- Yethiraj, A. and van Blaaderen, A. (2003) *Nature*, **421**, 513.
- Zandi, R. and Reguera, D. (2005) *Phys. Rev. E*, **72**, 021917.